Development and Testing of a High-Average Power, 94-GHz Gyroklystron

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Abstract—The development of a 10-kW average power, 94-GHz gyroklystron amplifier is described. This average power was obtained with 11% radio frequency (RF) duty factor and 92-kW peak power in the TE01 circular cavity mode. The instantaneous bandwidth was 420 MHz, and the efficiency was 33.5%. Low-duty-factor testing also yielded a peak power of as much as 115 kW with 600-MHz instantaneous bandwidth. This development effort was carried out over the past three years and represents record average power performance in an amplifier at this frequency.

Index Terms—Amplifier, gyro-klystron, gyrotwystron, millimeter wave, W-band.

I. INTRODUCTION

The development of gyrotwystron oscillators and amplifiers has been ongoing for more than three decades as a result of their unsurpassed capability in delivering high power at millimeter-wave frequencies [1], [2]. Gyrotron oscillators have found wide usage in such diverse applications as electron cyclotron resonance heating (ECRH) of fusion plasmas [3], industrial processing of ceramics [4], and dynamic nuclear polarization in conjunction with nuclear magnetic resonance (NMR) imaging [5].

Although the predominant emphasis of the gyrotron development community worldwide has been toward the development of gyrotron oscillators, gyrotron amplifiers have also received attention. This attention is the result of their potential as radio frequency (RF) drivers for linear colliders [6] and their application to high-power millimeter-wave radar. A variety of radar applications require the development of high-power millimeter-wave amplifiers such as gyroklystrons. Precision tracking of targets is often best accomplished with high-frequency radars [7]. The imaging of space objects, such as missiles, asteroids, space debris, and satellites, also requires the use of high-power millimeter-wave radars [8]–[13]. These radar applications have resulted in continuing interest in gyrotron amplifiers, including their many variants, such as the gyroklystron, gyrotwystron, and gyro-traveling wave tube (TWT). Of the many types of gyro-amplifiers that have been investigated in experimental low-average power devices, the most common type usually considered for radar application is the gyroklystron [14], [15].

Both this general need for millimeter-wave amplifiers for radar and the specific need of an amplifier for a planned Naval Research Laboratory (NRL) millimeter-wave radar have led to the development, over the past several years, of a high average power 94-GHz gyroklystron amplifier by a combined government, industry, and academic team [16]. It is the results of this team effort, reported on briefly in [16], that are described in detail in this paper. The paper is organized as follows. The issues addressed during the design procedure and the design are described in Section II, followed by presentation in Section III of the experimental test results. The experimental results are then discussed and compared with numerical codes in Section IV. Improvements to the codes that resulted from this careful comparison with experimental data are described. Finally, planned future directions and conclusions are presented in Section V.

II. GYROKLYSTRON DESIGN

The design of the gyroklystron amplifier was a major effort from both the physics and engineering points of view. The design procedure involved the use of many theoretical models and computer codes. These models and codes have been benchmarked over the past several years by a successful series of experimental low-duty-factor gyroklystron devices of narrow bandwidth at the University of Maryland, and of higher bandwidth at NRL [17]–[19]. Central to the suite of design codes are those codes that model the details of the beam–wave interaction in the gyroklystron circuit. There are two principal codes for this modeling activity, MAGYKL [20] (colloquially...
pronounced “magical”) and MAGY (“maggie”). The interaction circuit was designed primarily using a time-dependent version of the nonlinear code MAGYKL [17]. In this formalism, the cold-cavity RF electromagnetic fields are determined by a scattering matrix method [21], and then these fields are used for the integration of particle orbits within the circuit. The field amplitudes are self-consistently adjusted to represent the actual power transfer from the beam to the wave and thus achieve energy balance. However, the model for the fields in this code does not include modification of the longitudinal field profile in the cavities by the beam electric susceptibility, and in this sense, this code is not self-consistent. The fully self-consistent (in this sense) analysis code is MAGY [22], which was used primarily in the analysis of the experimental results reported herein, rather than in the original design of the device. Several other auxiliary interaction codes were also employed for particularly specialized computational tasks, including a linear theory analysis code, QPB [23], which was used to determine the stability of each cavity and drift section for the nominal operating parameters, and drift tube stability codes, as described below. A more detailed description of the theoretical tools and design methodology can be found in [24].

The electron gun design for this gyroklystron was of critical importance to the success of the development effort. The electron beam is formed by a 65-kV, 6-A magnetron injection gun that was designed to have very good beam quality, with perpendicular rms velocity spreads near 2% [25]. The electron gun was designed with a double anode configuration, namely, with a modulating anode (mod-anode) and an anode nominally at ground potential. The mod-anode is referenced to cathode potential, and beam cutoff condition may be achieved with a −2.5-kV mod-anode bias. The beam α is controlled by the voltage on the modulating anode (pulse-top voltage), which is nominally 17 kV above cathode potential, and by the superconducting coil that controls the magnetic field over the cathode.

The RF design goal for this gyroklystron development was 80-kW peak power at 12.5% duty factor for an average power of 10 kW. In addition, a goal of 600-MHz bandwidth (3 dB) at 94-GHz center frequency was the objective. A four-cavity circuit operating with circular TE_{011} cavity modes was chosen for the design. The circuit consists of a drive cavity (cavity 1), two bunching (idler) cavities (cavities 2 and 3), and an output cavity (cavity 4), all designed to operate in this mode at the fundamental of the cyclotron resonance. To determine the optimal circuit parameters, an extensive study was made of the tradeoffs in gain, power, and bandwidth that come through varying cavity and beam parameters. A complete description of this design study can be found in [26] and is, thus, not repeated here. For the optimized circuit parameters detailed in Table I, the predicted efficiency and peak output power versus frequency for a 65-kV, 6-A electron beam are shown in Fig. 1. As shown in the figure, efficiencies of up to 25% and bandwidths of approximately 600 MHz are expected for beam velocity ratios, \( \alpha = \frac{v_{beam}}{v_{c}} \), between \( \alpha = 1.5 \) and \( \alpha = 1.7 \) and an rms perpendicular velocity spread of 2.2%. The output power and efficiency are sensitive to relatively small changes in the beam velocity ratio.

Following the output cavity is a nonlinear up-taper to the collector, which also serves as the output waveguide. The shape of the taper was specially designed to minimize mode conversion while tapering up to a large radius in a short length [27]. The power is coupled out of the device through a chemical vapor deposition (CVD) diamond window, and the output mode is TE_{011}, which propagates in 1.2849 OD circular waveguide. Cold tests showed that for frequencies in the range of 93 GHz to 95 GHz, between 0.25% and 1% of the power incident on the window is reflected.

A. Cavity Electromechanical Design

From performance calculations with smooth wall cavities using the aforementioned design codes, the design cavity parameters are given in Table I. For optimum gyroklystron performance, the primary design requirement for all cavities is operation at the desired resonant frequency \( f_r \) and bandwidth or \( Q \). The input coupler has to satisfy the following additional requirements. The TE_{011}-mode should be excited with high mode purity and maximum power coupling. The input coupler design chosen was one in which the input waveguide excites...
the TE\(_{411}\)-mode in an outer coaxial cavity, and power is then coupled through four axial slots into the inner cylindrical cavity TE\(_{011}\)-mode. This type of input cavity design has been previously described [24], [28]. This input coupler was designed using both HFSS, a three-dimensional (3-D) finite-element electromagnetics code [29] and ARGUS, a 3-D finite-difference electromagnetics code [30]. The differing situations in which these two complementary codes were used in the design effort has been previously described [24].

For the operation of the amplifier at the full designed RF duty factor (12.5%), the drive-frequency–dependent average power load in the bunching cavities has a worst-case value of 400 W. Ohmic loading by itself is of course insufficient to achieve a loaded \( Q \) of 175. Furthermore, loading by the addition of ceramic rings internal to the cavities, which has been employed in low-duty-factor experimental devices at NRL and elsewhere, is too difficult to cool adequately at these high average power levels. Therefore, to meet this average power requirement for the operating mode, power is extracted from the bunching cavities to external loads. This yields a total cavity \( Q \) that is dominated by the external (or diffractive) \( Q \)-factor of these loads, and the heat load is dissipated externally.

After the design of the bunching cavities was finalized for the TE\(_{011}\) operating mode resonant frequency and \( Q \)-factor, the \( Q \)-factor of the cavities in other potentially competing modes (TE\(_{022}\), TE\(_{122}\), and TE\(_{212}\)) was also calculated and checked against the starting \( Q \)-factor of these modes. The starting \( Q \)-factor for a given parasitic mode is that \( Q \) that would cause the oscillation threshold in that mode to be reached at the design beam current. The loading schemes employed to achieve the requisite frequency and \( Q \) in the operating mode must in addition have \( Q \)-factors for the parasitic modes that are below the start-\( Q \) of those modes. The parasitic modes checked were the TE\(_{022}\) and TE\(_{212}\)-modes, with starting \( Q \)-factors of 800 and 425 and frequencies of 104.3 GHz and 88.9 GHz, respectively. The designed bunching cavities had a calculated \( Q \) of 192 and 68 in these modes, respectively, according to HFSS simulations. The electromagnetics code ARGUS [30] was also used in the design. ARGUS simulations yielded calculated \( Q \)-factors of 154 and 66, respectively, in excellent agreement with HFSS calculations. Consequently, the cavity design was believed to be stable to oscillations in these parasitic modes. Finally, the stabilization of the TE\(_{112}\)-mode was left to the drift-tube loading scheme (see below).

### B. Dielectric Probe Design and Cavity Testing

Even before beginning fabrication of the cavities, it was recognized that cold-test capability would be critical to the development of the proper fabrication procedure given the finalized cavity designs. The stagger tuning of the existing design would admit some variation in actual fabricated cavity resonant frequency and \( Q \) without significant loss of device performance (indeed, the design was optimized for a certain degree of insensitivity to exact frequencies and \( Q \)). However, the degree of cavity sensitivity to machined cavity diameter [118 MHz for a 5 \( \mu \)m (0.0002") change in diameter] necessitated the development of an accurate and precise cold-test procedure for quality control of fabricated cavities at each step of the assembly process. Space considerations precluded the usage of apertures in the cavities specifically for cold-test purposes. It was therefore decided to rely on TE\(_{01}\) axial transmission that was evanescently coupled through cavity irises to determine the cavity resonant frequency and \( Q \). This required the use of a high dynamic range W-band network analyzer and additional RF test equipment as follows. Marie converters were used to convert from the TE\(_{10}\)-mode in WR-10 rectangular waveguide to the TE\(_{01}\)-mode in a 0.2” diameter circular guide. To allow propagation of the TE\(_{01}\)-mode in waveguide diameters below air-loaded cutoff, diagnostic probes were fabricated using a dielectric-loaded circular waveguide with approximately the same diameter as the cavity iris. Each waveguide was connected to a nonlinear transition region to couple a TE\(_{01}\) wave in a 0.2” air loaded waveguide to a TE\(_{01}\) wave in 0.109” diameter Rexolite-loaded waveguide with minimal reflection and mode conversion [31].

Key issues that were addressed by the probe design were precision manufacture and the minimization of reflections in the transmission chain. At the small dimensions required for W-band operation, the return loss of the transition region is sensitive to the actual taper profiles, as well as the relative axial position of the wall and dielectric tapers. Tight tolerances are also required to permit “threading” of the probes through the brazed cavity stack. In addition, transmission profiles were found to be sensitive to small mismatches (∼25 dB) in the transmission chain. With precision manufacture of the dielectric probes, careful alignment, TE\(_{11}\)- and TE\(_{21}\)-mode filters, and Thru-Reflect-Line (TRL) calibration [32], [33] at the probe tips to minimize mismatch effects, consistent and reliable results have been obtained.

Resonant frequency measurements are found to be repeatable to within ∼40 MHz (better than 0.05%). \( Q \) measurements exhibit more variation, \( Q \_T \), typically, ∼150±20. Finally, the resonant frequencies measured using dielectric probes are generally within 50 MHz of frequencies measured with other methods, such as by transmission measurement through coupling holes in TE\(_{011}\)-mode test cavities used in NRL low-duty-factor experiments. Sample data for “circuit-ready” cavities are shown in Fig. 2. As a further indication of the accuracy possible using dielectric probes, Fig. 3 shows the measured and calculated char-
Fig. 3. Comparison of cavity transmission cold-test using dielectric probe and simulation of the same cavity using measured dimensions with HFSS electromagnetics code.

characteristics of an intermediate cavity. For the HFSS calculation, the measured mechanical cavity dimensions were used. Resistive wall losses were not included in this simulation.

C. Drift Tube Design

A critical issue in the design of the drift tubes is maintaining stability to spurious electromagnetic oscillations. Although the drift tube radius is deliberately made small enough to prevent TE$_{01}$-mode propagation at the operating frequency, parasitic oscillations can occur by one of two mechanisms. First, the radial wall transitions between the drift tube and the gyroklystron cavities on either side can create reflections, allowing the drift tube to act as an inadvertent resonant cavity to higher frequency modes. In the second mechanism, lower order modes, such as the TE$_{11}$ and TE$_{21}$, which can propagate at frequencies considerably below the TE$_{01}$ cutoff, can exhibit gyro–backward-wave instabilities. In order to achieve stability for both of these possible instability mechanisms, lossy-dielectric–lined drift tubes were employed.

Of the various arrangements of lossy dielectrics in the drift tubes that have been described in the literature [34], we decided to employ the structure of Fig. 4 because it is simple and highly effective [35]. This arrangement has been described previously [24].

The detailed design procedure is beyond the scope of the present paper; however, a short summary follows. The basic design process can be broken into three steps [24]. The first step is a study of the stability of a simple cylindrical drift tube without lossy dielectric, in which the troublesome modes are identified, the likely oscillation frequencies are computed, and the conditions for starting oscillation are examined. For a cavity-like oscillation, a maximum stable $Q$ can be determined through a linear start oscillation computation with test particles (in the same manner as was carried out for the cavities) [20], and for gyro–backward-wave oscillations, a critical length and growth rate can be computed with Vlasov theory [36]. The second step consists of selecting a suitable ceramic material for use as the lossy dielectric, and either determining the complex dielectric constant experimentally or relying on published values. The third step consists of modeling the dielectrically lined drift tube using an infinite-length, dielectrically loaded waveguide analysis to compute the modified dispersion curves, the attenuations per unit length, the modified oscillation frequencies, and the approximate quality factors of a finite length section of drift tube (if the ends were perfectly reflecting). The radial thickness of the dielectric is selected by repeating step three for a variety of dielectric thicknesses in an iterative fashion until acceptable values of attenuation and quality factor are achieved.

An instability in a drift tube occurs at a frequency near the intersection between the dispersion curve (a plot of frequency versus axial wavenumber) of an electromagnetic wave in the drift structure and the dispersion line for Doppler upshifted cyclotron waves supported by the electron beam. For the nominal gyroklystron operating point of 65-kV beam voltage with 36.8-kG magnetic field and a velocity ratio of 1.5, and using a slightly enlarged drift tube radius of 1.75 mm, analysis of the relevant dispersion curves indicates that a potential TE$_{11}$ backward-wave oscillation exists at about 76.7 GHz and a TE$_{11}$ cavity oscillation could occur near 119 GHz. Start-oscillation
analysis at a beam current of 6 A shows that the critical quality factor for the higher frequency oscillation is about 200, and the backward-wave growth rate of the low frequency oscillation is 14 dB/cm. A composite ceramic composed of 80% BeO and 20% SiC by weight was selected as the dielectric loading material. The complex dielectric constant at 94 GHz is about $\varepsilon + j\sigma$, where $\varepsilon$ is the complex dielectric constant of the dielectric, $\sigma$ is the inner radius of the dielectric, $b$ is the outer radius of the dielectric (assumed to be enclosed by metal beyond $r = b$), and $n$ is the azimuthal mode number. The dispersion relation must be solved numerically for the $k_z$ versus $\omega$ relationship to compute the intersections with the beam cyclotron dispersion. The infinitely long model is appropriate when the physical length of the drift tube is much longer than is the guide wavelength $2\pi/k_z$.

For a radial thickness $b - a$ of 0.8 mm, and retaining $a = 1.75$ mm, we obtain the dispersion diagrams shown in Fig. 5. Only the modes with an azimuthal number of unity are shown here, because for this system, all other azimuthal modes were found to be more stable. Potential oscillation frequencies are indicated by circles in the figure, and they consist of backward waves near 52, 56, and 76 GHz, and a forward wave at 119 GHz. The first two intersections are related to the original $TE_{11}$ and $TM_{11}$ dielectric-free solutions that have been almost completely drawn into the dielectric (to become hybrid $HE_{11}$- and $HM_{11}$-modes). (The $HE_{11}$- and $HM_{11}$-modes are sometimes also referred to as $HE_{11}$- and $EH_{11}$-modes in the literature). These solutions are heavily damped by the dielectric, with attenuations reaching 50 dB/cm in the 50–60-GHz range. The heavy damping plus the absence of field energy in the vacuum region completely eliminates any chance of oscillation. The third backward-wave and the forward-wave intersections lie on the $HE_{12}$-mode, which is related to the dielectric-free $TE_{12}$-mode. This hybrid mode has a modest portion of its field energy inside the dielectric, which leads to loading, and the remaining field structure in the vacuum region remains nearly identical to that of the dielectric-free $TE_{11}$-mode. The similarity of this particular field structure to a vacuum solution in an $r = a$ cylindrical waveguide allows the effect of lossy dielectric to be approximately modeled as a simple resistive wall loss, greatly simplifying the computation of the loaded start-oscillation conditions by test particle or Vlasov methods.

The results of calculations of attenuation rate versus frequency in the vicinity of the $HE_{12}$ backward-wave intersection point are shown in Fig. 6, for three different values of radial dielectric thickness. A radial thickness of 0.8 mm creates a peak attenuation of 20 dB/cm that is well positioned to load the 76-GHz backward-wave oscillation. A similar computation of the drift tube loading near the 119-GHz forward-wave intersection, performed for the 0.8-mm thick case, indicated that the loaded quality factor remains below 30 at frequencies up to 140 GHz. Because this value is well below the starting $Q$ of 200, stability to a cavity-like oscillation is easily ensured. As a result of these computations, a dielectric thickness of 0.8 mm was adopted for the experiments.

During this extensive design procedure for the gun, circuit, collector, and window, detailed thermal and mechanical stress analysis was also carried out. At the completion of the final electrical and mechanical design of the device reported on herein, the amplifier had been thoroughly analyzed from the physics, electrical, and mechanical engineering points of view.

III. EXPERIMENTAL TEST RESULTS

Device fabrication was carried out during mid-1997 to mid-1998. The resulting gyrokystron is shown in Fig. 7, prior to initial testing. A liquid-helium–cooled NbTi superconducting magnet was employed for all of the testing described in the
subsequent sections. Eventual system application of this gyroklystron will use closed-cycle-cooled superconducting magnets of either NbTi or high-temperature superconductor (HTS) [38] construction, but these magnets were not completed in time for the testing reported in this paper.

Experimental testing of this device was carried out from August 1998 through March 1999. The testing was divided into two parts, a low-duty-factor testing period (August 1998–October 1998) and a high-duty-factor testing period (October 1998–March 1999). Extensive measurements of the RF performance were made initially as part of the low-duty-factor testing period to characterize the RF peak-power performance of the amplifier.

Two different drivers were used in the experimental demonstration. For low-duty-factor testing (<0.2% RF duty factor), the drive power was supplied by an extended interaction oscillator (EIO), which is mechanically tunable from approximately 92.5 GHz to 95.5 GHz. The EIO produces up to 1.6-kW peak output power at pulse lengths up to 2 μs. For high-duty-factor testing, a coupled-cavity TWT (CPI VTW-6495) amplifier (TWTA) was used, together with a frequency synthesized solid-state source. This TWTA produces up to 80-W peak output power over the 93-GHz to 95-GHz frequency range at duty factors up to 100% (CW). The drive power is measured with a calibrated directional coupler at the gyroklystron input flange, and the frequency of the input and output RF signals were measured with a spectrum analyzer. Losses in the drive line between the source and gyroklystron input flange were found to be approximately 3 dB. A water load was positioned on the atmospheric side of the output vacuum window, and the temperature rise of the water for a given volumetric flow rate was used to measure the average RF power. In the experimental testing, the mod-anode voltage was provided by a compensated resistive divider network connected between the cathode and ground.

During the low-duty-factor testing, parametric studies of circuit operation with beam current and magnetic field variations were made at low RF and beam duty-factor to determine the best operating point for subsequent high-duty-factor operation. The solenoid consisted of a pair of main coils over the circuit, with each coil excited by a separate current, denoted $I_1$ and $I_2$. Generally, the current in the coil closest to the electron gun was fixed at $I_1 = 9.65$ A, and the second coil (current $I_2$) was used to vary the magnetic field over the circuit. $I_2$ was typically varied in the range of $I_2 = 8.9$ A to $I_2 = 9.2$ A, corresponding to a variation of between 400–800 G over the circuit. The field profiles obtained with these settings and the relative placement to the circuit is shown in Fig. 8. As shown in Fig. 9, the gyroklystron produced up to 118-kW peak output power and 29.5% efficiency in the TE_{311}-mode using a 66.7-kV, 6-A electron beam at 0.2% RF duty factor. At this operating point, the instantaneous full-width half-maximum (FWHM) bandwidth was 600 MHz and the gain was 24.7 dB. At higher magnetic fields, the bandwidth increases and the peak output power and gain decrease. This trend is predicted by the theoretical model, which shows that at lower fields, the input cavity operates in the negative beam-loading regime, which increases the $Q$ of this cavity and leads to lower bandwidth and higher gain and output power. As the magnetic field is increased, the cavity moves into the positive beam-loading regime, where the positive beam $Q$ decreases the overall $Q$ of the cavity, resulting in wider bandwidth and lower gain and output power. This behavior has also been documented in other detailed studies of bandwidth in gyroklystrons [39].

Also evident in Fig. 9 is the ripple in the peak output power across the frequency band. This effect has been determined to be caused by small reflections from the CVD diamond window.
The reflected wave, depending on the round-trip phase shift between the output cavity and window, can add either in phase or out of phase with the cavity fields. Because this phase shift is frequency dependent, the window reflection results in the modulation of the output power across the frequency band as a function of drive frequency. This effect was modeled with the MAGY code. The code showed that the spacing and amplitude of the output power across the band is consistent with reflections from the output window (see next section). The code also showed that the amplitude of the ripple is reduced at higher magnetic fields, as was observed experimentally. More detailed theoretical analyzes of the measured data will be presented in a companion paper [40].

Upon completion of the low-duty-factor demonstrations, the EIO was replaced with the 100-W coupled cavity TWTA driver for high-duty-factor tests. Because of input power limitations, a high-gain operating point was chosen for the high-duty-factor tests. During initial operation at high-duty-factor, 1.5-kW average power operation was achieved, and then upon increasing the duty factor to 2-kW average power, the output window developed a slow air-to-vacuum leak. The leak was patched, and subsequent testing of the window found anomalously high RF absorption in the diamond window. In addition, detailed simulations of the beam trajectories in the collector, including the trajectories of multiply scattered primary electrons, predicted that a small current could reach the output window. A minor flaw in the diamond window, if present, could also have contributed to the leak. Although all three factors (high RF absorption, possibility of flaw, and stray reflected primary electrons) could have contributed to the window failure, it was not determined conclusively which of the possibilities was the cause for failure. The remedy applied in the rebuilt tube was a new, lower loss diamond window, and an additional transverse beam-dump magnet near the output window. The dielectric loss tangent for the new window was measured, and the window was determined to have approximately 10 dB lower losses than does the first window.

The new window was installed, the gyroklystron reprocessed, and testing resumed.

Following a check that confirmed the same low-duty-factor operation with the reprocessed tube, the high-duty-factor testing was resumed. High average power operation was achieved in short order. Fig. 10 shows the measured peak and average output power versus drive frequency for a 66-kV, 4.15-A electron beam. The RF pulsewidth was 100 µs, and the pulse repetition frequency was 1.1 kHz, corresponding to 11% RF duty factor. The gyroklystron produced 10.1-kW average output power, corresponding to 92-kW peak output power and 33.5% efficiency. The measured FWHM bandwidth was 420 MHz. Fig. 11 shows the average output power and gain versus drive power. As shown in the figure, the maximum average output power of 10.1 kW and peak power of 92 kW is achieved for 53-W peak driver power, corresponding to 32-dB gain. High average power testing was halted at 11% duty factor rather than the designed 12.5% because of modulator limitations.

During high-duty-factor testing, the bandwidth of 420 MHz was obtained, compared with the 600-MHz bandwidth obtained in low-duty-factor testing. This inability to reach 600 MHz in high-duty-factor testing resulted from the limitations of the test set and the TWTA driver. In order to achieve the 10-kW average power with the limited power available from the TWTA driver, it was necessary to operate the gyroklystron at the lower magnetic field parameters that gave narrower bandwidth and higher gain. It is believed that with a higher power driver, or with a few decibels additional gain, the 600-MHz bandwidth achieved in low-duty-factor testing would be achieved at high duty factor. A second gyroklystron presently under development is expected to remedy this situation.

Fig. 12 shows the peak output power versus drive frequency for RF pulsewidths of 10 µs, 40 µs, and 70 µs. For each curve, the beam current and beam voltage pulsewidths were held fixed at 90 µs, and the pulse repetition frequency was 1 kHz. As
Fig. 11. Measured average output power and gain for a 66 kV, 4.15 A electron beam and 11% rf duty factor. The drive frequency was held fixed at 93.8 GHz.

Fig. 12. Peak output power for RF beam pulse widths of 10 μs (open circle), 40 μs (open square), and 70 μs (x). For each curve, the beam pulse width, pulse repetition frequency, beam current, beam voltage, and all other parameters were held fixed.

seen in the figure, for RF duty factors ranging from 1% (10-μs pulsewidth) to 7% (70-μs pulsewidth), the differences in peak output power over the operating bandwidth were negligible. The measured results show that for duty factors in this range (1–7%), there were no significant changes in the circuit or circuit performance resulting from thermal effects. In addition, circuit heating from RF losses, as measured by independent cooling loops, was at the nominal (calculated) values up to the full duty factor tested (11%).

Parametric studies were also performed at high duty factor. By increasing the magnetic field, bandwidths up to 600 MHz could be obtained with reduced peak and average output powers, typically, near 50 kW and 5 kW, respectively. The theoretically predicted trend of a reduction in the ripple on the gain frequency response at the higher magnetic fields was also observed at high duty factor. In addition, calorimetric measurements were made for the window, the interaction circuit, and other cooled components of the device and were found to be close to the theoretically predicted values of losses.

In early April 1999, testing was completed and the gyroklystron was removed from the test stand at Communications and Power Industries (CPI). The gyroklystron has subsequently been shipped to NRL for delivery to the NRL Radar Division.

IV. ANALYSIS AND DISCUSSION

In order to gain insights into the performance of the gyroklystron, a series of simulations with the time-dependent, self-consistent code MAGY [22] have been performed. In contrast to the model used in MAGYKL [20], MAGY does not assume a priori the axial RF field profiles; rather, the field profiles are allowed to evolve self-consistently in response to beam–wave interactions. This capability is crucial for accurately predicting some phenomena in gyroklystrons. More detailed discussion on self-consistent effects in gyroklystrons is presented in a separate paper in this issue [40]. In addition, a model for window reflections is incorporated into MAGY, thus permitting evaluation of the experimentally observed ripples in frequency response.

A detailed comparison of experimental data and numerical simulations has been carried out for two operating beam currents, 4 A and 6 A, and for a range of magnetic field values. Both frequency response curves (output power versus drive frequency) and drive curves (output power versus input power) have been compared. A representative sample of these numerous comparisons is presented here.

Shown in Fig. 13 are the comparison between experimental data and MAGY simulation of the frequency response (a) and drive curve (b) for a beam current of 4 A, at given magnetic field values determined by superconducting solenoid currents \( I_1 = 9.65 \) A and \( I_2 = 9.0 \) A. The cold-cavity resonance frequencies and quality factors used in the simulations are as shown in Table II. These values agree with cold-test values within the experimental uncertainty of the cold-test measurement technique at the time the cold tests were performed [31]. The input powers shown were measured at the input flange and include the 3-dB known loss between this flange and the Te_{0q} cavity.

Because the actual beam \( \alpha_d \) was not measured directly, it was determined via a combination of the output cavity Te_{0q} start-oscillation data and EGUN scaling of the magnetron injection gun design [25]. The methodology for determining the beam velocity ratio \( \alpha_d \) is as follows. A series of MAGY simulations are first performed for a given beam current and magnetic field profile to determine the threshold \( \alpha_d \) at which the output cavity begins to oscillate. Simulations with small variations below the start oscillation \( \alpha_d \) are then performed to determine the beam \( \alpha_d \) that best matches a set of experimental bandwidth and drive curves obtained just below the start-oscillation threshold. This provides us with the base \( \alpha_d \) from which beam \( \alpha_d \) for other sets of data can be determined, with EGUN scaling based on the relative change of the experimentally known gun voltages or magnetic field. In the EGUN scaling and in MAGY simulations, the
Fig. 13. Comparison of MAGY code and experimental data for 4 A beam current, and at magnetic field given by \( I_2 = 9.0 \) A for both the frequency response (a) and the drive curve (b). Solid curves with diamond markers is code prediction; filled circles are experimentally measured data.

### TABLE II

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magnetic field profiles used are obtained with the magnetic field code EFFI [41], using actual experimental settings of the superconducting coils. The EFFI calculated and measured axial magnetic field profiles have been compared and found to be in excellent agreement. In addition, a window reflection of 1% (\(-20\)-dB return loss) from the CVD diamond window, based on cold-test measurements, has also been incorporated into the MAGY simulations shown in Fig. 13. Fig. 14 shows a similar comparison between experimental data and the MAGY code predictions for the case of a 6-A beam current. The excellent agreement between the experimental data and the simulations provides compelling evidence that the observed frequency response ripple can indeed be attributed to the small window reflection. In addition,
it is clear that the overall performance of the gyroklystron is well modeled by the MAGY code.

As a practical note, although this ripple in frequency response is nonideal, it is of relatively minor consequence in the radar application. Plotted on a logarithmic scale, it amounts to on the order of ±0.5-dB variation from the no-reflection curves; this level of variation is straightforward to remove in the radar signal processing. Nevertheless, one of the goals for subsequent devices currently under development is a reduction of this frequency response ripple.

V. CONCLUSION

In summary, a four-cavity, TE_{011}, mode, high-average-power gyroklystron amplifier was designed, built successfully, and tested. The successful development, in a single generation, of this high average power device is a testament to both the validity of the design procedure and the quality of the physical models and corresponding computer codes that were employed in this effort. At low duty factor, the amplifier produced up to 118-kW peak output power at 600-MHz bandwidth. At high duty factor,
10.1-kW average output power and 420-MHz bandwidth were achieved at 11% RF duty factor with an efficiency of 33.5%. At different magnetic field settings, the device produced 600-MHz instantaneous bandwidth with 5-kW average power and 50-kW peak power. These results represent world-record–setting performance for an amplifier at this frequency.

Future work will be centered on improving certain aspects of the device performance. Because of the limited availability and high cost of high-duty-factor drivers with tens of watts of power at millimeter-wave frequencies, future work will be focused on improving device gain, in anticipation of being able to saturate with solid-state drivers in subsequent devices. A subsequent amplifier currently in construction will comprise five cavities to improve the gain of the device by more than 10 dB. Future work will also consider bandwidth improvement by replacing the output cavity with a traveling-wave output section, a configuration commonly referred to as a gyrotron amplifier. A low-duty-factor prototype of a W-band gyrotron has demonstrated 50-kW peak output power and 925-MHz FWHM bandwidth, and calculations of projected performance with the present electron gun design are for 1.4-GHz bandwidth at 80-kW peak power level [42]. For bandwidths greater than 2 GHz, gyro-TWT amplifiers are under investigation by a number of researchers [43]–[45], and there is ample reason to be optimistic about the future development of high average power versions of gyro-TWT’s in the millimeter-wave band.

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